

EARTH ORBITING DARK MATTER AS A POSSIBLE  
EXPLANATION OF THE FLYBY ANOMALIES

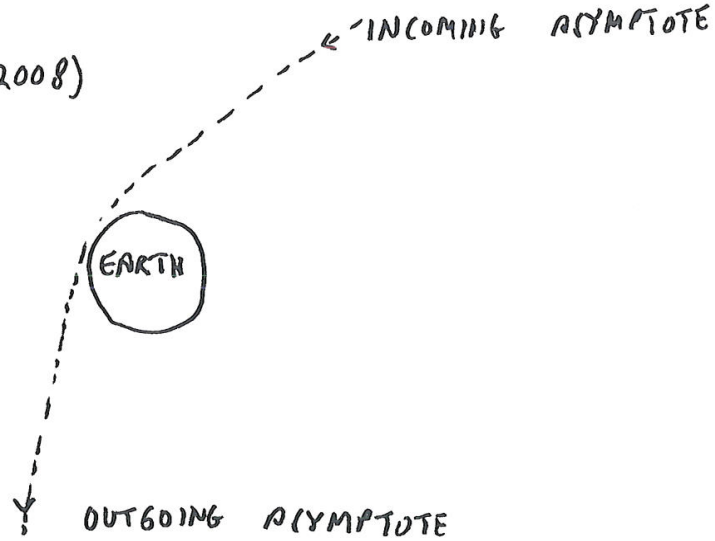
Stephen L. Adler  
Institute for Advanced Study

- EXPERIMENTAL DATA Anderson et al PRL 100  
091102 (2008)
- POSSIBLE EXPLANATIONS
- DARK MATTER SCATTERING ? S.L.A. Phys. Rev. D 79  
023505 (2009)
- BOUNDING THE MASS OF EARTH  
ORBITING DARK MATTER S.L.A. J. Phys. A Math. Theor.  
41 (2008) 412002
- DARK MATTER MODELING S.L.A. arXiv: 0908.2919
- SPACECRAFT CALORIMETRY AS  
A TEST S.L.A. arXiv: 0910.1564  
"white paper" for decadal  
review on space science
- OUTLOOK

EXPERIMENTAL DATA

Anderson et. al.  
PRL 100, 091102 (2008)

Anderson + Nieto  
Physic Today Oct 2009  
pp 76-77



OUTGOING VELOCITY EXTRAPOLATED FROM INCOMING VELOCITY DOES NOT AGREE WITH MEASURED VALUE

PARAMETER	GALILEO I	GALILEO II	NEAR	CASSINI	ROSETTA	MENESGER
DATE	12/8/90	12/8/92	1/23/98	8/18/99	3/4/05	8/2/05
$\Delta V_{\infty}$ ( $\frac{mm}{s}$ )	3.92	-9.6	13.96	-2	1.80	0.02
$\leftarrow V_{\infty}$ ( $\frac{mm}{s}$ )	0.3	1.0	0.01	1	0.03	0.01

ANDERSON et al FIT (EMPIRICAL - NO THEORETICAL BASIS)

$$\frac{\Delta V_{\infty}}{V_{\infty}} = \frac{1}{2} \frac{\Delta E}{E} = K (\cos \delta_i - \cos \delta_o)$$

$\delta_{i,o}$  = INCOMING, OUTGOING DECLINATION  
↗  
LATITUDE ANALOG IN  
CELESTIAL COORDINATE  
SYSTEM

$$K = \frac{2\omega_E R_E}{c} = 3.099 \times 10^{-6}$$

$\omega_E$  = EARTH ANGULAR ROTATION VELOCITY

$R_E$  = EARTH RADIUS  $\approx 6,371$  km

- DECLINATION DEPENDENCE INDICATES THAT MAXIMAL SYMMETRY OF ANY MECHANISM CAN BE AXIAL SYMMETRY AROUND EARTH ROTATION AXIS
- SPHERICAL SYMMETRY CANNOT GIVE A DECLINATION DEPENDENCE

-4-

## MAGNITUDE AS AN ACCELERATION

NEAR FLYBY:  $\Delta V_{\infty} = 13.46 \text{ mm/s} \sim 1.3 \cdot 10^{-2} \text{ m/s}$   
OVER  $\Delta T \sim 4 \text{ HOURS} \sim 10^4 \text{ s}$   
 $\Rightarrow$  ACCELERATION  $\sim 10^{-6} \text{ m/s}^2$

SOME COMPARISONS:

PIONEER ANOMALOUS ACCELERATION  $\approx 8.74 \cdot 10^{-10} \text{ m/s}^2$

MOND CUTOFF ACCELERATION  $\sim 10^{-10} \text{ m/s}^2$

COSMOLOGICAL SCALE  $H_0 c \sim 2.3 \cdot 10^{-18} \text{ s}^{-1} \cdot 3 \cdot 10^8 \frac{\text{m}}{\text{s}} \sim 7 \cdot 10^{-10} \frac{\text{m}}{\text{s}^2}$

FLYBY ANOMALY, INTERPRETED AS AN ACCELERATION,  
IS 3 ORDERS OF MAGNITUDE LARGER THAN THESE

## POSSIBLE EXPLANATIONS

### FOUR POSSIBILITIES:

- EFFECT IS AN ARTIFACT - (SOME ESSENTIAL CONVENTIONAL PHYSICS HAS BEEN OMITTED FROM THE ORBITAL CALCULATION)
  - THERMAL EFFECTS (WHICH MAY EXPLAIN MOST OF THE PIONEER ANOMALY) ARE TOO SMALL TO EXPLAIN FLYBY (BEAN WORKSHOP REPORTS)
  - OTHER POSSIBILITIES UNDER STUDY
- NEW ELECTROMAGNETIC PHYSICS
  - QED TESTED TO HIGH PRECISION
  - PVLAS - DIDN'T FIND AXIONS, BUT SHOWED VACUUM LINEAR TO  $1$  IN  $10^{18}$  IN 2.5 TESLA FIELD
- NEW GRAVITATIONAL PHYSICS
  - PPN DEVIATIONS FROM GENERAL RELATIVITY AT LEAST A FACTOR OF 100 TOO SMALL (SEE MEMO ON S.L.A. WEB PAGE TALKS + MEMOS SECTION)

## EXPLANATIONS - CONTINUED

SO A GRAVITATIONAL EXPLANATION MUST BE OUTSIDE THE PPN FRAMEWORK OF METRIC THEORIES THAT OBEY THE EQUIVALENCE PRINCIPLE

- BUT EQUIV. PRIN. TESTED TO HIGH PRECISION

• MOND OR COSMOLOGY: ACCELERATIONS TOO SMALL

• DARK MATTER

COSMOLOGY  $\Rightarrow$  ONLY  $\sim 4\%$  OF THE MASS-ENERGY DENSITY OF THE UNIVERSE IS BARYONIC MATTER

$\sim 23\%$  IS GRAVITATIONALLY ATTRACTIVE

"DARK MATTER"

$\sim 73\%$  IS GRAVITATIONALLY REPULSIVE

"DARK ENERGY"

LITTLE IS KNOWN ABOUT DARK MATTER:

BOSONIC OR FERMIONIC?

SELF-ANNIHILATING?

MASS (G) ?

NON-GRAVITATIONAL INTERACTIONS ?

## DARK MATTER SCATTERING?

DARK MATTER CAN BE GRAVITATIONALLY BOUND  
ON DIFFERENT SCALES

- GALACTIC HALO DARK MATTER

MASS DENSITY  $\rho_m \sim 0.3 \text{ GeV}/c^2 \text{ cm}^{-3}$

- SOLAR SYSTEM-BOUND DARK MATTER?

$$\rho_m < 10^5 \text{ GeV}/c^2 \text{ cm}^{-3}$$

FROM STUDY OF PLANETARY ORBITS

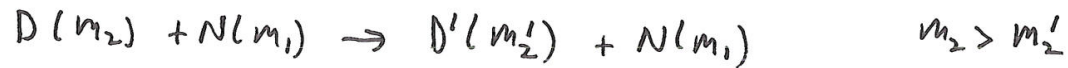
- EARTH-BOUND DARK MATTER?

CAN SCATTERING FROM EARTH-BOUND DARK  
MATTER EXPLAIN THE FLYBY ANOMALIES?

SCATTERING - CONTINUED

- ELASTIC SCATTERING - GIVES NORMAL DRAG - VELOCITY DECREASE

- INELASTIC EXOTHERMIC SCATTERING



CAN GIVE A VELOCITY INCREASE - WILL GIVE DETAILED FORMULAS

- TO GIVE OBSERVED VELOCITY ANOMALIES, NEED

ELASTIC  $\sigma \bar{\rho}_m \geq 10^{-16} \text{ GeV/c}^2 \text{ cm}^{-1}$   
↑  
(CROSS SECTION)

INELASTIC

$$\sigma \bar{\rho}_m \geq 10^{-20} \text{ GeV/c}^2 \text{ cm}^{-1}$$

$$(\text{INELASTIC} / \text{ELASTIC} \sim c/v_{\text{ORBITAL}} \sim 10^4)$$

$$\sigma \sim 10^{-33} \text{ cm}^2 \rightarrow \bar{\rho}_m \text{ ELASTIC} \sim 10^{17} \text{ GeV/c}^2 \text{ cm}^{-3}; \bar{\rho}_m \text{ INELASTIC} \sim 10^{13} \text{ GeV/c}^2 \text{ cm}^{-3}$$



PLACING DIRECT LIMITS ON THE MASS OF  
EARTH-BOUND DARK MATTER

CAN SET A DIRECT LIMIT ON THE TOTAL  
EARTH-BOUND DARK MATTER MASS LYING BETWEEN  
MOON ORBIT RADIUS  $\sim 384,000$  km

LAGEOS GEODETIC  
SATELLITE ORBIT RADIUS  $\sim 12,300$  km

FOR A CIRCULAR ORBIT, MEASURING RADIUS  $R$   
AND PERIOD  $T$  GIVES  $GM$  OF ATTRACTING BODY,

$$GM = \frac{4\pi^2 R^3}{T^2}$$

• LAGEOS TRACKING GIVES  $GM_{\oplus}$

EARTH MASS  $M_{\oplus}$  INCLUDES DARK MATTER  
WITHIN THE LAGEOS ORBIT

DIRECT LIMITS - LUNAR ORBITER, ERUS TRACKING,  
LUNAR LASAR RANGING

- LUNAR ORBITERS GIVE  $GM_m$
- MORE ACCURATE  $M_m$  FROM TRACKING ERUS ASTEROID GIVES

$$R_{\oplus l_m} = \frac{GM_{\oplus} + G\Delta M_{\oplus}}{GM_m} \quad \leftarrow \begin{array}{l} \text{POSSIBLE DARK MATTER} \\ \text{EFFECT ON ERUS} \end{array}$$
$$= \frac{GM_{\oplus}}{GM_m} (1 + \delta) \quad \delta = \frac{\Delta M_{\oplus}}{M_{\oplus}}$$

- LUNAR LASAR RANGING GIVES COMBINED  $GM_{TOT}$  OF EARTH-MOON SYSTEM:

$$GM_{TOT} = GM_{\oplus} + GM_m + GM_{dm}$$

↑  
MASS OF DARK MATTER  
LYING BETWEEN MOON AND  
LARGER ORBITS

DIRECT LIMITS - FINAL FORMULA

COMBINING FORMULAS

$$\begin{aligned}
GM_{TOT} - GM_{\oplus} - \frac{GM_{\oplus}}{R_{\oplus/m}} &\cong GM_{dm} + \frac{M_m}{M_{\oplus}} GM_{\oplus} \\
&\cong GM_{dm} + 0.0123 GM_{\oplus} > GM_{dm} \\
&= GM_{dm} [1 + O(0.01)] \quad \text{IF } \Delta M_{\oplus} \sim M_{dm}
\end{aligned}$$

NUMERICAL (ALL CONVERTED TO BARYCENTRIC DYNAMICAL TIME)  
(SIYVA Tsvytchen, JPL)

LAGEOS  $\rightarrow GM_{\oplus} = 398,600.4356 \pm 0.0008 \text{ km}^3 \text{ s}^{-2}$

LUNAR RANGING  $\rightarrow GM_{\text{combined}} = 403,503.2357 \pm 0.0019 \text{ km}^3 \text{ s}^{-2}$

EROS  $\rightarrow R_{\oplus/m} = 81.300570 \pm 0.000005$

+ LAGEOS  $GM_{\oplus} \rightarrow GM_m = 4902.8000 \pm 0.0003 \text{ km}^3 \text{ s}^{-2}$

DIRECT LIMITS- RESULTS

COMBINING THESE GIVES

$$GM_{dm} \leq (0.0001 \pm 0.0016) \text{ km}^3 \text{ s}^{-2} = (0.3 \pm 4) \times 10^{-9} GM_{\oplus}$$

→  
DOMINANT ERROR FROM LUNAR LASER RANGING

$$\text{so } M_{dm} \lesssim 4 \cdot 10^{-9} M_{\oplus}$$

IF THIS BOUND WERE ATTAINED, AND THE MASS WERE UNIFORMLY DISTRIBUTED BELOW THE MOON'S ORBIT, THE DENSITY WOULD BE

$$\rho_m \sim 6 \cdot 10^{10} \text{ GeV/c}^2 \text{ cm}^{-3} \Rightarrow \text{GALACTIC HALO DENSITY AND LIMIT ON SUN-BOUND DARK MATTER}$$

IF CONFINED TO A NARROW BAND,  $\rho_m$  FOR EARTH-BOUND DARK MATTER COULD BE EVEN HIGHER: FOR A SHELL OF RADIUS R AND THICKNESS D,

$$\rho_m \lesssim \frac{2 \cdot 10^{10} \text{ GeV/c}^2 \text{ cm}^{-3} (389,000 \text{ km})^3}{R^2 D}$$

## DARK MATTER MODELING FOR FLYBY

SOME FLYBYS GAIN VELOCITY (ANOMALOUS DRAG)

SOME FLYBYS LOSE VELOCITY (NORMAL DRAG)

MODEL ASSUMES TWO SPECIES OF DARK MATTER  
ORBITING EARTH

- INELASTIC, EXOTHERMIC SCATTER

$$D(m_2) + N(m_1) \rightarrow D'(m_2') + N(m_1) \quad m_2 > m_2'$$

CAN GIVE VELOCITY GAIN IF  $N(m_1)$  SCATTERS FORWARD

- ELASTIC SCATTER

$$D''(m_2'') + N(m_1) \rightarrow D''(m_2'') + N(m_1)$$

GIVES NORMAL  
DRAG

VELOCITY CHANGE FORMULAS

$\vec{u}_1$  = VELOCITY OF INITIAL NUCLEON (IN SPACECRAFT)

$\vec{u}_2$  = VELOCITY OF INCIDENT DARK MATTER

$\langle \delta \vec{v} \rangle$  = CROSS-SECTION AVERAGED NUCLEON VELOCITY CHANGE  
(ASSUME NO AZIMUTHAL ANGLE DEPENDENCE)

INELASTIC

$$\langle \delta \vec{v} \rangle = \frac{\vec{u}_1 - \vec{u}_2}{|\vec{u}_1 - \vec{u}_2|} A_I \quad A_I = \left[ \frac{2 (m_2 - m_2') m_2'}{m_1 (m_1 + m_2')} \right]^{1/2} c \langle \cos \theta \rangle_I$$

ELASTIC

$$\langle \delta \vec{v} \rangle = -(\vec{u}_1 - \vec{u}_2) A_E \quad A_E = \left( \frac{m_2''}{m_1 + m_2''} \right) [1 - \langle \cos \theta \rangle_E]$$

WILL LUMP CROSS SECTION X DENSITY INTO THE  
PARAMETERS  $A_I$ ,  $A_E$

## CHANGE IN OUTGOING SPACECRAFT VELOCITY

FORCE = VELOCITY CHANGE IN SINGLE SCATTER • NUMBER OF SCATTERS PER UNIT TIME

$$\delta \vec{F} = \int d^3 u_2 \langle \delta \vec{v} \rangle |\vec{u}_1 - \vec{u}_2| \sigma(\vec{x}, \vec{u}_2)$$

$$\vec{u}_1 = d\vec{x} / dt$$

INTEGRATING  $\int \vec{F} \cdot d\vec{x}$  ALONG TRAJECTORY

$$\delta \frac{1}{2} (\vec{V}_f^2 - \vec{V}_i^2) = \vec{V}_f \cdot \delta \vec{V}_f = \int_{t_i}^{t_f} dt (d\vec{x}/dt) \cdot \delta \vec{F}$$

$$= \int_{t_i}^{t_f} dt \int d^3 u_2 \frac{d\vec{x}}{dt} \cdot \langle \delta \vec{v} \rangle |\vec{u}_1 - \vec{u}_2| \sigma(\vec{x}, \vec{u}_2)$$

CROSS SECTION AND SCATTERING-ANGLE AVERAGES

$W =$  CENTER OF MASS SCATTERING ENERGY

$$\frac{W}{(m_1 + m_2) c^2} \approx 1 + \frac{m_1 m_2}{2(m_1 + m_2)^2} \frac{(\vec{u}_1 - \vec{u}_2)^2}{c^2}$$

FOR  $m_2 \leq m_1$  AND NON-RELATIVISTIC  $\vec{u}_1, \vec{u}_2$ ,  $W \approx 1$

$\Rightarrow$  CROSS SECTION DOMINATED BY LOWEST PARTIAL WAVES

ELASTIC

S - WAVE DOMINATED  $\sigma \approx \sigma_{el} = \text{CONSTANT}$

$$\langle \cos\theta \rangle_E = 0$$

$$1 - \langle \cos\theta \rangle_E = 1$$

INELASTIC

ENTRANCE CHANNEL  
MOMENTUM

$$k = \frac{m_1 m_2}{m_1 + m_2} |\vec{u}_1 - \vec{u}_2|$$

$$\frac{d\sigma}{d\Omega} = \frac{A_{\text{inel}}}{4\pi} k^{-1} + B_{\text{inel}} \frac{3}{4\pi} \cos\theta + \dots \quad \text{NEAR THRESHOLD}$$

$$\sigma \approx A_{\text{inel}} k^{-1} \quad \langle \cos\theta \rangle = B_{\text{inel}} / (A_{\text{inel}} k^{-1})$$

$$\sigma \langle \cos\theta \rangle_I = B_{\text{inel}} = \text{CONSTANT}$$

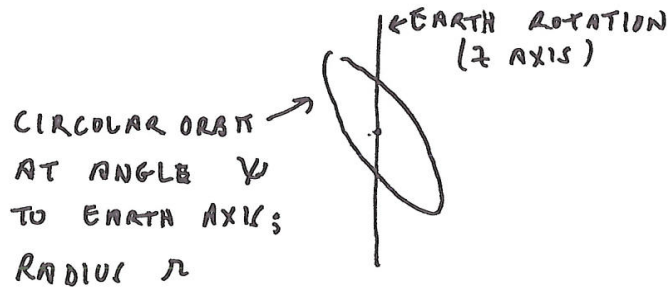


DARK MATTER ORBITAL GEOMETRY

ASSUME AXIAL SYMMETRY AROUND EARTH ROTATION AXIS

SIMPLEST MODEL - DISK IN EQUATORIAL PLANE - DOES NOT WORK - "NEAR" GETS SMALLEST VELOCITY CHANGE

NEXT SIMPLEST MODEL -



$E, L_z$  CONSERVED, BUT ORBIT WILL PRECESS BECAUSE OF EARTH QUADRUPOLE - WILL FILL SHELL OBTAINED BY AVERAGING ORBIT OVER AZIMUTHAL ANGLE AROUND EARTH AXIS



← GET THIS

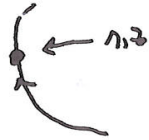
$$\text{DENSITY} \propto \frac{1}{r} \underbrace{\frac{1}{\sqrt{(r \sin \psi)^2 + z^2}}}_{\text{JACOBIAN}}$$

$w(r, \psi)$

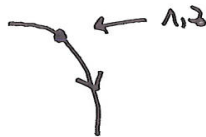
↑  
WEIGHTING FUNCTION

DARK MATTER VELOCITY  $\vec{u}_2$

FOR GIVEN  $\lambda, \psi, z$ , THERE ARE TWO  $\vec{u}_2$  VALUES, ONE FOR AN UP-GOING SEGMENT



AND ONE FOR A DOWN-GOING SEGMENT



$$\vec{u}_2 = \sqrt{\frac{GM_{\oplus}}{a}} (c_{\pm} \hat{u}_{||} + d_{\pm} \hat{u}_{\perp})$$

$\hat{u}_{||}$  PARALLEL TO EQUATORIAL PLANE  
 $\hat{u}_{\perp}$  PERPENDICULAR TO  $\hat{v}$  AND  $\hat{u}_{||}$

$$c_{\pm} = \frac{\lambda \cos \psi}{\sqrt{\lambda^2 - z^2}}$$

$$d_{\pm} = \pm \frac{\sqrt{\lambda^2 \sin^2 \psi - z^2}}{\sqrt{\lambda^2 - z^2}}$$

$$c_{\pm}^2 + d_{\pm}^2 = 1$$

SIMPLE CHOICE OF  $w(r, \psi)$

SINGLE TILT ANGLE  $\psi_i, \psi_e$  FOR INELASTIC, ELASTIC SHELLS

GAUSSIAN  $r$  PROFILE

$$w_i(r, \psi) = K_i e^{-\frac{(r-R_i)^2}{D_i^2}} f(\psi - \psi_i) \quad \text{INELASTIC}$$

$$w_e(r, \psi) = K_e e^{-\frac{(r-R_e)^2}{D_e^2}} f(\psi - \psi_e) \quad \text{ELASTIC}$$

$$\begin{aligned} N = \text{TOTAL NUMBER OF PARTICLES IN SHELL} &= \int d^3x \int d^3u_2 \rho(\vec{x}, \vec{u}_2) \\ &= 4\pi^{5/2} K_l D_l \quad l = i, e \end{aligned}$$

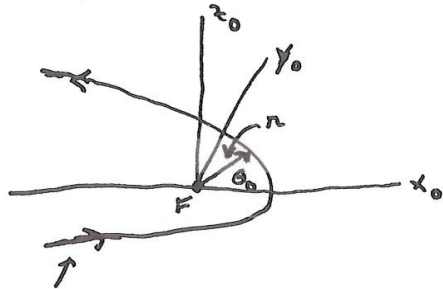
DENSITY TIMES CROSS SECTION  
PARAMETERS:

$$\rho_i \equiv N_I B_{incl} K_i$$

$$\rho_e \equiv N_E \sigma_{el} K_e$$

# FLYBY ORBITAL GEOMETRY

WORK IN FLYBY ORBIT PLANE



FLYBY HYPERBOLIC ORBIT

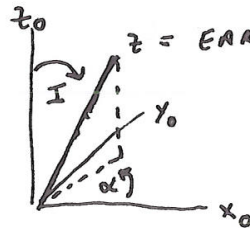
$$\lambda = \frac{\mu}{1 + e \cos \theta_0}$$

$$\vec{r} = (x_0, y_0, 0)$$

$$x_0 = \lambda \cos \theta_0$$

$$y_0 = \lambda \sin \theta_0$$

EARTH AXIS AS SEEN FROM FLYBY ORBIT PLANE



z = EARTH ROTATION AXIS  
POLAR ANGLES I, alpha

- FIRST, EXPRESS DARK MATTER ORBIT ON EARTH-FIXED AXES
- THEN, REWRITE ON FLYBY PLANE AXES TO GET

$$z, \hat{n}_{\parallel}, \hat{n}_{\perp}$$

FLYBY PARAMETERS ON FLYBY PLANE AXES  
(AND INTRINSIC PARAMETERS  $\rho, e$ )

	<u>GL-I</u>	<u>GL-II</u>	<u>NEAR</u>	<u>CASSINI</u>	<u>ROSETTA</u>	<u>MESSENGER</u>
$I(^{\circ})$	142.4	138.7	108.0	25.4	144.9	133.1
$\alpha(^{\circ})$	-45.1	-147.4	-55.1	-158.4	-53.1	0.0
$\rho$ (km)	25,473	22,155	19,448	51,689	19,262	20,563
$e$	2.473	2.319	1.814	5.851	1.312	1.360

$$e = 1 + \frac{2 V_{\infty}^2}{V_f^2 - V_{\infty}^2}$$

$$\rho = \frac{q G M_{\oplus}}{V_{\infty}^2} \left[ \left( \frac{V_{\infty}^2}{V_f^2 - V_{\infty}^2} \right)^2 + \frac{V_{\infty}^2}{V_f^2 - V_{\infty}^2} \right]$$

$\alpha$  REQUIRES ATTENTION TO  $\pm$  IN ARC COSINE ; ABOVE VALUES REPRODUCE THE  $\delta_i, \delta_0, \alpha_i - \alpha_0$  VALUES IN ANDERSON ET AL PRL

NUMERICAL RESULTS

MINIMIZE LIKELIHOOD FUNCTION  $\chi^2 = \sum_{k=1}^6 (\delta V_{ksth} - \delta V_{kSA})^2 / \sigma_{kSA}^2$

$\delta V_{ksth} = \rho_i \delta V_{kSi} + \rho_e \delta V_{kSe}$  LINEAR IN  $\rho_i, \rho_e$  ;  $\chi^2$  QUADRATIC

⇒ MINIMIZATION WITH RESPECT TO  $\rho_{i,e}$  CAN BE DONE ALGEBRAICALLY

REDUCED PARAMETER SPACE FOR NUMERICAL SEARCH TO 6 PARAMETERS:  $\psi_i, R_i, D_i$   $\psi_e, R_e, D_e$

- WITH SMOOTHED JACOBIAN, DO COARSE MESH SURVEY TO IDENTIFY REGIONS WITH LOW  $\chi^2$  TO USE AS STARTS FOR MINIMIZATION
- WITH ORIGINAL JACOBIAN, AND ADAPTIVE INTEGRATION, DO MINIMIZATION USING CERN PROGRAM "MINUIT"

ILLUSTRATIVE GOOD FIT (2d FROM PREPRINT)

	$\chi^2$	GL-I	GL-II	NEAR	CAGINI	ROSETTA	MESSENGER
$\delta V_A$ (mm/s)		3.92	-4.6	13.96	-2	1.80	0.02
$\sigma_A$ (km/s)		0.3	1.0	0.01	1	0.03	0.01
FIT 2d	0.51	3.90	-4.6	13.96	-2.7	1.80	0.020

↓  
PARAMETERS

$10^6 \times \rho_i$ (km)	$10^2 \times \rho_e$ (km)	$\psi_i$ (rad)	$\psi_e$ (rad)	$R_i$ (km)	$D_i$ (km)	$R_e$ (km)	$D_e$ (km)
1.000	0.288	1.372	0.3902	34520	3030	29370	6678

OVERFITTING?

- WITH ONLY INELASTIC, CANNOT FIT NEAR + MESSENGER  $\chi^2 = 0.61 \times 10^5$
- WITH FULL MODEL, FITTING TO 5 OF 6 FLYBOYS GIVES QUALITATIVELY REASONABLE VALUE FOR 6<sup>th</sup>:

3.71	-4.6	16.03	-2.7	1.62	0.12
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⇒ NOT JUST FITTING A WIGGLY CURVE THROUGH DATA POINTS

IMPLICATION OF DARK MATTE MASS BOUND

LUNAR LASER RANGING  $\Rightarrow 6 (M_{\oplus} + M_m + M_{DM})$

LAGEOS  $\Rightarrow 6 M_{\oplus}$

LUNAR ORBITER  $\Rightarrow 6 M_m$   
EROS TRACKING

COMBINING  $\Rightarrow M_{DM} < 4 \times 10^{-9} M_{\oplus} \sim 1.4 \times 10^{23} \text{ GeV}/c^2$

FIT 2d:  $\rho_i D_i = 0.00304 \text{ km}^2$

$\rho_e D_e = 19.2 \text{ km}^2$

USE  $M_e \equiv m_2 N_e = 4\pi^{5/2} \rho_e D_e m_1 / \sigma_{el}$

$M_i \equiv m_2 N_i = 4\pi^{5/2} \rho_i D_i m_1 / \beta_{inel}$

$M_e \leq 4 \times 10^{-9} M_{\oplus} \Rightarrow \sigma_{el} \geq 9.4 \times 10^{-31} \text{ cm}^2$

$M_i \leq 4 \times 10^{-9} M_{\oplus} \Rightarrow \beta_{inel} \geq 1.5 \times 10^{-34} \text{ cm}^2$



TEMPERATURE CHANGE

FLYBY VELOCITY CHANGE USED  $\langle \delta \vec{v} \rangle = \frac{\int_0^\pi d\theta \sin\theta \delta \vec{v} |f(\theta)|^2}{\int_0^\pi d\theta \sin\theta |f(\theta)|^2}$

FLOUCTUATION

$\langle (\delta \vec{v} - \langle \delta \vec{v} \rangle)^2 \rangle$

GIVES A TEMPERATURE INCREASE

$\delta T_{\text{single collision}} = \frac{m_1}{2 k_B} \langle (\delta \vec{v} - \langle \delta \vec{v} \rangle)^2 \rangle$

$\frac{dT}{dt} = \int d^3 u_2 \delta T_{s.c.} \sigma(|\hat{u}_1 - \hat{u}_2|) n(\vec{x}, \vec{u}_2)$

$T_f - T_i = \int_{t_i}^{t_f} dt \frac{dT}{dt}$

SIMPLE MODIFICATION TO PROGRAMS

DEFINE:  $\frac{R_{\text{inel}}}{B_{\text{inel}}} \equiv R_{\text{inel}} m_2 c$        $S_{\text{inel}} = R_{\text{inel}} \left[ \frac{2(m_2' - m_2)m_2'}{m_2} \right]^{1/2}$

GET THEN:  $\frac{\Delta T}{^\circ K} = \left( \frac{m_2 c^2}{\text{MeV}} \right) [z_e + S_{\text{inel}} z_i]$

FIND:  $z_e \sim (0.3 - 0.9) \times 10^{-5}$        $z_i \sim (0.2 - 0.7) \times 10^4$

KEY THINGS NEEDED FROM FLYBY REANALYSIS

- REAL EFFECT (NEW PHYSICS) OR ARTIFACT?
- $\int U_f$  VALUES
- $\sigma$  VALUES  
 $\sigma = .01$  MESSENGER, NEAR  
 $\sigma = .03$  ROSETTA  
GIVES THEM A VERY BIG WEIGHT  
IN  $\chi^2$
- FLYBY ORBITAL PLANE PARAMETERS  
AND EARTH AXIS POLAR ANGLES  $I$   
AZIMUTH  $\alpha$

ALSO LOOK FOR

- RESULTS FROM DIRECT DARK MATTER SEARCHES
- LHC SEARCH FOR SUPERSYMMETRY