(1) In class, we discussed the metric

$$ds^2 = -2e^q dv du + g_{AB}(dx^A + c^A dv)(dx^B + c^B dv),$$

where q is independent of u at v = 0, but otherwise the functions q, g_{AB} , and c^A depend on all coordinates u, v, x^A . As discussed in class, this is a canonical form for the metric near a null hypersurface Y that is swept out by a family of orthogonal null geodesics from a codimension 2 spacelike submanifold W. W is the hypersurface u = v = 0 and Y is the hypersurface v = 0.

Calculate R_{uu} along Y, that is at v = 0. As explained in class, this step leads to Raychaudhuri's equation.

(2) Consider the metric

$$ds^2 = (t^2 - 1) \left(-dt^2 + d\vec{x}^2 \right),$$

where $\vec{x} = (x^1, \dots, x^{D-1}).$

In this spacetime, consider the codimension 2 spacelike hypersurface W defined by $t = t_0$, $|\vec{x}| = R$, with constants t_0, R . What is the condition for W to be a trapped surface?

(3) In a spacetime M, let S be a spacelike hypersurface (dimension D-1 if M has dimension D) and let $Q \subset S$ be a manifold with boundary, also of dimension D-1; let ∂Q be the boundary of Q. For example, in Minkowski space, Q might be a closed ball and then its boundary ∂Q is a sphere. As usual, we let $J^+(Q)$ be the causal future of Q and $J^+(\partial Q)$ be the causal future of ∂Q . The boundaries of these sets are $\partial J^+(Q)$ and $\partial J^+(\partial Q)$.

Show that any point in $\partial J^+(Q)$ that is not in Q itself (in other words, any point that is strictly to the future of Q) is in $\partial J^+(\partial Q)$. This fact will be useful in discussing black holes.

1